# Using the MCNP6.2 Correlated Fission Multiplicity Models, CGMF and FREYA

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## Introduction

- MCNP is a general purpose Monte Carlo transport code
  - Many choices must be made by users to run the code properly for their applications
  - Example: What kind of variance reduction techniques are needed for my shielding application?

## • Applications include:

- Criticality safety
- Nuclear safeguards and non-proliferation
- Nuclear energy
- Nuclear threat reduction and response
- Radiation detection and measurement
- Radiation health protection
- Stockpile stewardship

## **Motivation**

What results do I want to obtain from my Monte Carlo calculation?

- Fixed-source
  - current, flux, charge and energy deposition, dose
- Criticality
  - k<sub>eff</sub>, current, flux, reaction rates, power distributions, burnup

#### All of the above can be computed using default MCNP fission treatment

- Fixed-source
  - time-coincident detector pulses (e.g. multiplicity counters)

#### Need fission multiplicity treatment - not default in MCNP

## **Motivation**

- Warhead Measurement Campaign (WMC) meant to passively and actively measure nuclear warheads for treaty verification
  - New measurements of neutron and photon coincidence data of shielded special nuclear materials (SNM)
  - At the time, the transport simulation tools available were limited in their ability to fully predict WMC-like measurements
  - This was due to the type of nuclear fission data available
  - To address these shortcomings, more detailed behavior of nuclear fission physics was needed
  - Making the transport simulations more predictive in SNM detection applications

#### • Key Issues

- Average nuclear data quantities are insufficient
- Need better ways to compare to experiment

# Background

- MCNP6 default nuclear fission physics
  - Average photon production for each collision
  - Average neutron production for each fission
  - Average energy spectra for neutrons and gamma rays
  - Isotropic angular emission
  - No correlations!
- Applications
  - Shielding: current, flux, energy deposition, dose
  - <u>Subcritical / Critical Systems</u>: k<sub>eff</sub>, flux, reaction rates
  - <u>Reactor Physics</u>: k<sub>eff</sub>, current, flux, power distributions, burnup
  - <u>Radiation Detection</u>: charge and energy deposition, pulse-height spectra, bulk counting rates





# Background

- MCNP6.2 nuclear fission physics available
  - Multiplicity distribution of gamma rays for each fission
  - Multiplicity distribution of neutrons for each fission
  - Multiplicity dependent energy spectra (energy correlations)
  - Angular emission from fission fragments (angular correlations)
  - Full correlations!

## • Applications

- In addition to MCNP6 default applications ...
- <u>Subcritical Systems</u>: singles, doubles, etc. counting rates, leakage multiplication, probability of initiation/extinction
- <u>Reactor Physics</u>: higher-order power distribution fluctuations
- <u>Radiation Detection</u>: n-n, n-γ, γ-γ time coincidence





# Background MCNP Default Behavior

- In nature, a fission event will emit a number of neutrons and gamma rays with some probability distribution, P(E, v)
- In MCNP, by default
  - The average neutrons emitted is used,  $\overline{v}(E)$
  - Bounded integer sampling scheme:
    - If  $\overline{v} = 2.2$ ,
    - Then, P(v = 2) = 80% and, P(v = 3) = 20%
  - Preserves expected value

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Induced Fission Neutron Multiplicities for Plutonium-239



# Background FMULT Input Card

- Original FMULT options within MCNP turns on neutron multiplicity sampling and allows the user to,
  - Modify spontaneous fission average multiplicity and yield rate
  - Change Watt energy spectrum parameters for spontaneous fission
  - Provide Gaussian FWHM width for spontaneous and induced fission multiplicity distributions
  - Select a sampling algorithm and data source
- Does not handle fission gamma ray emission
- Each neutron emitted,
  - Direction is isotropic and independently sampled
  - Energy is sampled independently from the same energy distribution (uncorrelated)

See MCNP6 User's Manual, Los Alamos National Laboratory, LA-CP-14-00745 (2014).

#### MCNP output file, print table 38

		-			
1fission	multip	licity dat	a.		
zaid	width	watt1	watt2	vield	sfnu
				1	
90232	1.079	.800000	4.00000	6.00E-08	2.140
92232	1.079	.892204	3.72278	1.30E+00	1.710
92233	1.041	.854803	4.03210	8.60E-04	1.760
92234	1.079	.771241	4.92449	5.02E-03	1.810
92235	1.072	.774713	4.85231	2.99E-04	1.860
92236	1.079	.735166	5.35746	5.49E-03	1.910
92238	1.230	.648318	6.81057	1.36E-02	0.048
93237	1.079	.833438	4.24147	1.14E-04	2.050
94236	0.000	.000000	0.00000	0.00E+00	0.080
94238	1.115	.847833	4.16933	2.59E+03	0.056
94239	1.140	.885247	3.80269	2.18E-02	2.160
94240	1.109	.794930	4.68927	1.02E+03	0.063
94241	1.079	.842472	4.15150	5.00E-02	2.250
94242	1.069	.819150	4.36668	1.72E+03	0.068
95241	1.079	.933020	3.46195	1.18E+00	3.220
* 96242	1.053	.887353	3.89176	2.10E+07	0.021
96244	1.036	.902523	3.72033	1.08E+07	0.015
96246	0.000	.000000	0.00000	0.00E+00	0.015
96248	0.000	.000000	0.00000	0.00E+00	0.007
97249	1.079	.891281	3.79405	1.00E+05	3.400
98246	0.000	.000000	0.00000	0.00E+00	0.001
98250	0.000	.000000	0.00000	0.00E+00	0.004
98252	1.207	1.180000	1.03419	2.34E+12	0.002
98254	0.000	.000000	0.00000	0.00E+00	0.000
100257	0.000	.000000	0.00000	0.00E+00	0.021
102252	0.000	.000000	0.00000	0.00E+00	0.057
* = use	d in pro	oblem.			

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# Background LLNL Fission Library

- In MCNP6.1 (and 6.1.1), the LLNL Fission Library is one of two (lowenergy) event generators:
  - Spontaneous, neutron-induced and photo-fission for most fissionable systems
  - MCNP6.2 version produces same results from previous versions
  - Input card: FMULT method=5
  - Now includes FREYA 2.0
- The LLNL Fission Library includes more tabulated and fitted data used for lesser known isotopes FREYA can't presently handle
- If the LLNL Fission Library cannot handle a particular isotope, the FMULT default
   parameters are used instead
   Default MCNPX treatment
   LLNL fission library treatment



# Background FREYA – LBNL/LLNL

- The FREYA is developed by LBNL/LLNL
- In MCNP6, it is accessible through LLNL Fission Package with different FMULT card option
  - Input card: **FMULT method=6**
- Spontaneous fission: <sup>238</sup>U, <sup>238</sup>Pu, <sup>240</sup>Pu, <sup>242</sup>Pu, <sup>244</sup>Cm, and <sup>252</sup>Cf
- Neutron-induced fission: <sup>233</sup>U, <sup>235</sup>U, <sup>238</sup>U, <sup>239</sup>Pu, and <sup>241</sup>Pu
- If FREYA cannot handle particular isotope, LLNL Fission Library is used
- FREYA uses a Monte Carlo Weisskopf approach
  - Neutrons emitted by sampling from Weisskopf spectrum
  - After neutrons are done emitting, gamma rays are emitted from residual energy
- Computationally more efficient than Monte Carlo Hauser-Feshbach (CGMF)



rrrr

BERKE



Fragment Mass (amu)

# Background CGMF – LANL

- CGMF is a superset of CGM with an added fission reaction capability
  - Input card: FMULT method=7
- Fission fragments are sampled from a joint probability distribution function of mass (A), charge (Z) and total kinetic energy (TKE)
- Uses Hauser-Feshbach statistical theory of nuclear reactions treating neutron / photon competition de-excitation of fission fragments
- Monte Carlo is used to sample each step in the de-excitation process
- Spontaneous fission: <sup>240</sup>Pu, <sup>242</sup>Pu, and <sup>252</sup>Cf
- Neutron-induced fission: <sup>235</sup>U, <sup>238</sup>U, and <sup>239</sup>Pu
- If CGMF cannot handle a particular isotope, the LLNL Fission Library is used instead





# Background FREYA & CGMF



# Background FREYA & CGMF



# Background Other Monte Carlo Codes

#### • MCNPX-PoliMi:

- Implemented some correlated models (IPOL(1)=1, IPOL(1)=10) for Cf-252 systems
- Integrated CGMF and FREYA into local research version

#### MCNPX (and now MCNP6)

• Implemented LLNL Fission Library

## • TRIPOLI (CEA)

LLNL Fission Library & FREYA integration supported by J. Verbeke

## • MORET (IRSN)

• LLNL Fission Library & FREYA

## • GEANT

• LLNL Fission Library & FREYA (only locally integrated, not in the open community)

## **Guidance with Numerical Results**

### Brief summary of options

- Default: no neutron multiplicity, no gamma-ray multiplicity\*, uncorrelated\*\*
- FMULT: neutron multiplicity, no gamma-ray multiplicity\*, uncorrelated\*\*
- LLNL: neutron multiplicity,
  - gamma-ray multiplicity, uncorrelated\*\*

gamma-ray multiplicity, correlated

gamma-ray multiplicity, correlated

- FREYA: neutron multiplicity,
- CGMF: neutron multiplicity,

n(1.0273 MeV)+<sup>239</sup>Pu fission reaction

Model	$\bar{\nu}_n$	$ar{\chi}_n$	$ar{ u}_{oldsymbol{\gamma}}$	$ar{\chi}_{\gamma}$
Default	3.0126(1)	2.139(1)	*	*
FMULT	3.012(1)	2.137(1)	*	*
LLNL	3.014(1)	2.036(1)	7.307(3)	0.8985(3)
FREYA	3.012(1)	2.153(1)	6.876(2)	1.0098(4)
CGMF	3.048(1)	2.033(1)	7.905(3)	0.9293(3)

\*gamma-rays can be present, but may not be fission reaction specific \*\*neutrons and gamma-rays are uncorrelated from each other



## **Guidance with Numerical Results**

### • Differential Experiment Validation

- Using differential measurements to validate event-byevent predictions of CGMF/FREYA
- CNEC student at LANL working on modeling detailed detectors in NEUANCE (M. Pinilla)
- Using PTRAC and DRiFT to model stilbene

## • To make sure CGMF & FREYA are a reflection of reality







## **Guidance with Numerical Results**

## • Simplified MCNP model of NEUANCE

- 21 stilbene detectors in array
- <sup>252</sup>Cf spontaneous fission source





- Using mcnptools (included with MCNP6.2 release) to analyze PTRAC results
- Need to add detector response using DRiFT (in progress)

# **Preview of Works in Progress**

- University of Michigan differential measurements of angular correlations
- Priority is to compare against experimental measurements
- Follow-up of 2014 NSE paper by S.A. Pozzi *et al*.
- Submitted an abstract to IRRMA X meeting in Chicago, IL, July 9-13
- Transport and post-processing code comparisons
  - MCNP6 / DRiFT
  - MCNP6 / MPPost
  - MCNPX-PoliMi / MPPost
  - MCNPX-PoliMi / DRiFT





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## **Preview of Works in Progress**



MCNP6.2 simulation of correlated counts



- Time-stamp differences are small
- Angular differences are significant!
  - Lots of cross-talk between detectors at small angles

## **Preview of Works in Progress**

#### Correlated counts



- Changing pulse height thresholds changes:
  - Overall count rates
  - Removes some of the cross-talk effect
- Time stamp differences change w.r.t angle



## **Conclusions & Future Work**

#### Conclusions

- MCNP6.2 manual is updated with new information on the fission multiplicity treatment options within **FMULT** and how they work
- When correlated fission multiplicity models are needed, MCNP6.2 now contains two such models:
  - FREYA from LBNL/LLNL
  - CGMF from LANL
- New post-processing capabilities distributed with MCNP6.2 for users in many application areas (see mcnptools for PTRAC processing)

#### **Future Work**

- Improvements to FREYA and CGMF
  - More isotopes/energies, photofission, time-dependent gamma-ray emission, etc.
  - Computational speed improvements (CGMF)
- Validation more simulation vs. experiment

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